## Meteor detection on spectrograms by means of computer vision

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## Automated detection - intro

## Antonio Martínez at the IMC2022:

## Narrow pass-band filtering technique for radio meteor automatic detection

Establishing a reliable method for detecting meteors in BRAMS project observations has proven to be an elusive issue. In this paper, a narrow filter-based technique is applied to a series of random spectrograms. The results of this proof-of-concept are promising

Fourteen years of Speciab 5 minutes records (VVS beacon) Felix Verbelen: 17 years, including .wav files



## Automated detection - intro

Similar image analysis approach:

WGN, the Journal of the IMO 47:2 (2019) 55 Automated Spectrogram Analysis for Meteor Head Echoes

C. Powell



## Typical cases



Resolution: time frequency

5 min 300 s 200 Hz 600 pixels 400 pixels

Minimum 'duration' of a reflection: 1 pixel = 0.5 s

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950 Hz

## **Reflections on lightning**







## **Classification** vs Object Detection

## It is a cat

## Specialist

## Machine















## Classification vs Object Detection

# Found a meteorSpecialistMachine+True+False



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False

True

## Preparing: convert to B&W



Color <> amplitude of the sound signal

## Preparing: getting rid of the 1 minute lines



- 1 x 3 pixels interpolation
- Time calibration

## **Frequency summation**



## • Carrier removal



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## **Time summation**



5a 202209210635CS sumf(t).png



# Real time: trigger on amplitude threshold

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## **Frequency summation**



## • Carrier removal

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## **Time summation**





# Real time: trigger on amplitude threshold

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## **Computer Vision**

**OpenCV** (**Open** Source **C**omputer **V**ision Library) is an open source **computer vision** and **machine learning** software library.

The library has more than 2500 optimized algorithms, which includes a comprehensive set of both classic and state-of-the-art computer vision and machine learning algorithms. These algorithms can be used to detect and recognize faces, identify objects, etc



## Thresholding



optimum threshold

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## Removal of 'straight' lines

Carrier (horizontal line) removal worked well Possibility to remove plane streaks (not horizontal)?

Hough transform (patent 1962): identification of lines in an image

Bubble chamber: charged particles create ionisation track





## Removal of 'straight' lines

Hough transform on spectrogram  $\rightarrow$  creates far too many lines Required: edge detection first (Canny, John - 1986)



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## Removal of 'straight' lines

- Do not detect 'vertical' lines = edges of meteors !
- Hough transforms allows 'gaps' (interruptions) in lines and mimimum length



## **Connected** components

Connect pixels with one out of the 8 neighbours exceeding a threshold



## **Connected** components



- Threshold 10, minArea 6
  - Threshold 20, minArea 6, Fillfactor 0.1, h/w 0.4
  - Threshold 20, minAread 6, Fillfactor 0.2, h/w 0.4

## **Connected** components



Threshold 15, minArea 6, Fillfactor 0.2, h/w 0.4, Blur 3x2, linear regression slope, correlation, spread

## Tuning the model

- Few cases → not representative
- Tune (manually) the parameters on a sufficient large training set 2023 Febr 16 – 18 (3 days, 493 meteors)
- Annotate true positives and false negatives
- Potential additional criteria: linear regression of the object pixels



## **Performance metrics**

- Recall (sensitivity) = True Pos / (True Pos + False Neg) = 88.8 % (False Neg = non meteors misidentified as meteors)
- Any pattern in the False Neg (Feb 18) ?



Yes, 'far' from the central frequency Second pass: eliminate for  $|\Delta y| > 60$ , removes 17 False Neg, looses 3 True Pos

New recall = 96.8 %

## **Performance metrics**

 Precision = True Pos / (True Pos + False Pos) = 92.6 % (False Pos = meteors not detected)



## Tuning the model

- 'Final' set of parameters
  - Threshold
  - Canny edge (2)
  - Hough lines (4)
  - Blurring
  - -h/w
  - Correlation coefficient
  - Area
  - Covariance (2)
  - Δf

## • Grid search / random search

## Advantage of automated detection

- Consistency
- No `positivity bias'
- Analyse vast volumes
- No special hardware required

 This study: only tuned for non-shower meteors (analyse Jan 5 – Apr 15, September)



## **Opportunities of automated detection**

- Counts weighted by duration
- Counts weighted by (audio) power



 $dB_i = 10 \log P_i \quad P_i = 10^{\frac{dB_i}{10}} P_{comp} = \sum_i P_i$  $dB_{comp} = 10 \log P_{comp}$ 

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## **Opportunities of automated detection**

- Power distribution: equivalent of magnitude distribution
- 'Radio' population index: discern sporadics from stream meteors
- With stable frequency or known carrier frequency: Doppler shift of centroid → distribution of velocity drifts, stream vs sporadics

## **Environment and packages**



## Other setups

- Source code available 'as is' to try yourself (tuned on mainly sporadic activity, probably performs less well on streams)
- A Felix Verbelen recording, partially tuned:



## Thanks to / acknowledgments

- Antonio Martínez Picar
- **Pierre Terrier**
- **Cis Verbeeck**
- Felix Verbelen
- Aegide Steyaert



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